

# $g - 2$ and New Physics

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Based on **arXiv:1104.1769** by **G.-C. Cho, K. Hagiwara, DN**  
and **Yu Matsumoto**



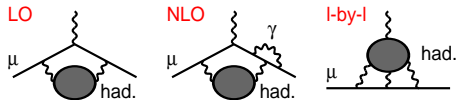
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# Introduction: Standard Model prediction for muon $g - 2$

<b>QED</b> contribution	11 658 471.808 (0.015) $\times 10^{-10}$	Kinoshita & Nio, Aoyama et al
<b>EW</b> contribution	15.4 (0.2) $\times 10^{-10}$	Czarnecki et al
<b>Hadronic</b> contribution		
<b>LO</b> hadronic	694.9 (4.3) $\times 10^{-10}$	HLMNT11
<b>NLO</b> hadronic	-9.8 (0.1) $\times 10^{-10}$	HLMNT11
<b>light-by-light</b>	10.5 (2.6) $\times 10^{-10}$	Prades, de Rafael & Vainshtein
<b>Theory TOTAL</b>	<b>11 659 182.8 (4.9) <math>\times 10^{-10}</math></b>	
<b>Experiment</b>	<b>11 659 208.9 (6.3) <math>\times 10^{-10}</math></b>	world avg
<b>Exp – Theory</b>	<b>26.1 (8.0) <math>\times 10^{-10}</math></b>	<b>3.3 <math>\sigma</math> discrepancy</b>

(Numbers taken from HLMNT11, arXiv:1105.3149)

n.b.: hadronic contributions:



## Muon $g - 2$ :

- ✓ Powerful probe for New Physics at TeV scale
- ✓  $\sim 3\sigma$  deviation between exp and theory (SM)  
 $\implies$  Signal of new physics?
- ✓ Possible signal from EW sector of New Physics  
 $\implies$  What about other EW precision tests?

Now, the question is:

Suppose that a New Physics is responsible for the muon  $g - 2$  anomaly. **Is it possible for the New Physics to be compatible with the final LEP EW data?**

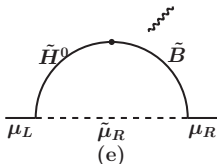
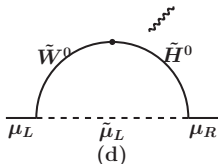
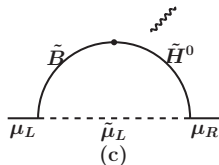
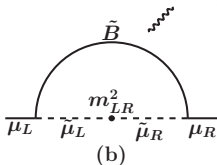
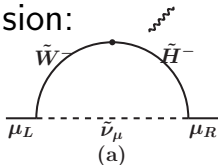
— **Important question to study BEFORE the LHC**

In this talk I take the MSSM as an example of new physics, since it is one of the most attractive models.

# SUSY Contributions to Muon $g - 2$

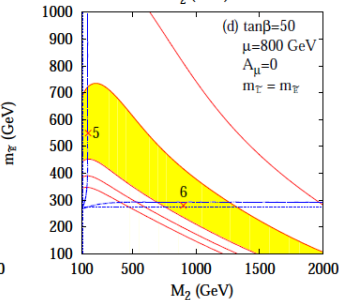
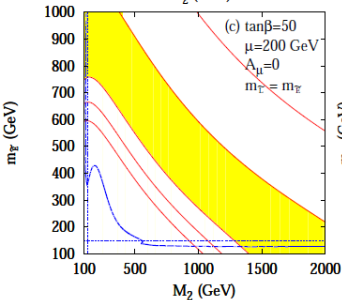
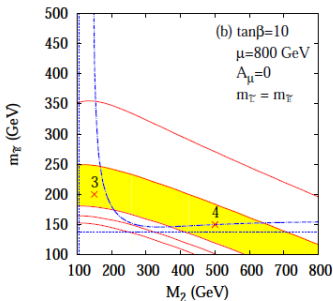
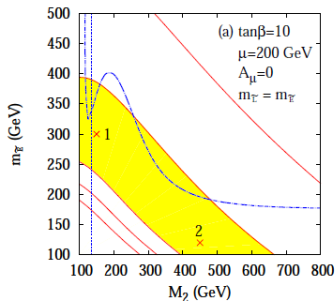
Suppose that the  $\sim 3\sigma$  deviation is due to SUSY...

Leading SUSY contributions in the  $m_Z/m_{\text{SUSY}}$  expansion:



In most cases, the  $\tilde{\chi}^\pm - \tilde{\nu}$  diagram (a) and/or the  $\tilde{B} - \tilde{\mu}_{L/R}$  diagram (b) dominate. (Lopez-Nanopoulos-Wang, Chattopadhyay-Nath, Moroi, ...)

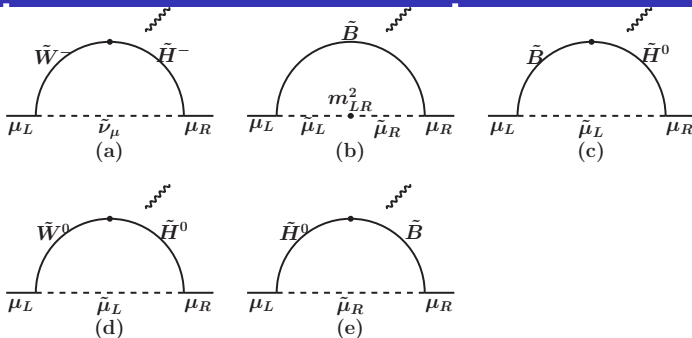
# MSSM Contributions to Muon $g - 2$



x-axis:  $M_2$   
 (gaugino mass)

y-axis:  $m_{\tilde{l}}$   
 (slepton mass)

# Muon $g - 2$ at MSSM sample points



No.	$\tan \beta$	$\mu$	$M_2$	$m_{\tilde{E}}$	(a)	(b)	(c)	(d)	(e)	(a)-(e)	total
1	10	200	150	300	29.6	1.1	0.7	-2.9	-1.3	27.2	25.0
2	10	200	450	120	27.5	8.8	3.3	-7.1	-6.7	25.9	25.9
3	10	800	150	200	14.3	16.2	0.6	-2.7	-1.3	27.1	27.1
4	10	800	500	150	6.9	21.3	1.0	-2.5	-2.1	24.7	24.3
5	50	800	150	550	26.9	2.4	0.5	-2.6	-1.0	26.3	26.0
6	50	800	900	280	18.0	18.0	2.5	-5.9	-5.1	27.7	27.6

The chargino diagram (a) and/or the Bino-smuon $_{L,R}$  diagram (b) dominate in all the sample points.

# Selected SUSY models and muon $g - 2$

## Selected SUSY Models

	$\tan\beta$	$\mu$	$m_{\tilde{\mu}_L}$	$m_{\tilde{\mu}_R}$	$A_\mu$	$M_1$	$M_2$
SG 1 (mSUGRA, $\tan\beta = 10$ )	10	396	181	116	-445	103	193
SG 2 (mSUGRA, high $\tan\beta$ )	50	762	585	465	-145	277	510
GM 1 (Gauge Med., high $\tan\beta$ )	42	504	441	214	25	181	339
GM 2 (Gauge Med., $\tan\beta \sim 10$ )	15	300	257	120	-39	169	327
MM1 (Mirage Med., $\alpha > 0$ )	10	430	188	255	-465	170	258
MM2 (Mirage Med., $\alpha < 0$ )	10	-572	253	108	245	-99	-248
MM3 (Mirage Med., $M_2 < M_1$ )	10	534	200	237	509	224	173

## Muon $g - 2$ in the Selected SUSY Models

	(a)	(b)	(c)	(d)	(e)	(a)-(e)	total	pull
SG 1	25.7	21.5	1.5	-5.2	-5.4	38.1	37.6	1.2
SG 2	20.0	4.8	1.0	-3.4	-2.8	19.5	19.4	-1.0
GM1	34.6	11.7	1.4	-5.3	-9.2	33.2	33.0	0.7
GM2	27.1	10.6	1.6	-5.0	-9.0	25.3	24.8	-0.3
MM1	19.4	7.2	1.4	-4.5	-1.9	21.7	21.7	-0.7
MM2	13.2	18.8	0.7	-2.7	-4.2	25.8	24.7	-0.4
MM3	19.6	7.9	1.1	-3.8	-1.8	23.0	23.1	-0.5



# Introduction to EW Precision Study

- **LEP-I** ('89 - '95): The  $Z$ -boson properties studied in great detail using 17 millions of  $Z$  boson decays. (Final report appeared in 2005: [hep-ex/0509008](http://hep-ex/0509008))
- To confront the EW precision data with theory, the  **$S, T, U$  parameters** are useful ([Peskin+Takeuchi](#)).

$$\gamma \text{---} \bullet \text{---} \gamma = i e^2 \Pi_{\text{QQ}} g^{\mu\nu} + \dots$$

$$Z \text{---} \bullet \text{---} \gamma = i \frac{e^2}{c s} (\Pi_{3Q} - s^2 \Pi_{\text{QQ}}) g^{\mu\nu} + \dots$$

$$Z \text{---} \bullet \text{---} Z = i \frac{e^2}{c^2 s^2} (\Pi_{33} - 2s^2 \Pi_{3Q} + s^4 \Pi_{\text{QQ}}) g^{\mu\nu} + \dots$$

$$W \text{---} \bullet \text{---} W = i \frac{e^2}{s^2} \Pi_{11} g^{\mu\nu} + \dots$$

$$\alpha S \equiv 4e^2 [\Pi'_{33}(0) - \Pi'_{3Q}(0)],$$

$$\alpha T \equiv \frac{e^2}{s^2 c^2 m_Z^2} [\Pi_{11}(0) - \Pi_{33}(0)],$$

$$\alpha U \equiv 4e^2 [\Pi'_{11}(0) - \Pi'_{33}(0)].$$

In this talk, we use an improved version,  **$S_Z, T_Z$  and  $M_W$**  ([Hagiwara, Haidt, Kim & Matsumoto](#)).

# $S_Z$ and $T_Z$ (1)

We define the 'bar charges' as

$$\begin{aligned}\bar{e}^2(q^2) &\equiv \hat{e}^2(\mu) \left[ 1 - \bar{\Pi}_{T,\gamma}^{\gamma\gamma}(q^2) \right], & \bar{s}^2(q^2) &\equiv \hat{s}^2(\mu) \left[ 1 + \frac{\hat{c}(\mu)}{\hat{s}(\mu)} \bar{\Pi}_{T,\gamma}^{\gamma Z}(q^2) \right], \\ \bar{g}_Z^2(q^2) &\equiv \hat{g}_Z^2(\mu) \left[ 1 - \bar{\Pi}_{T,Z}^{ZZ}(q^2) \right], & \bar{g}_W^2(q^2) &\equiv \hat{g}_W^2(\mu) \left[ 1 - \bar{\Pi}_{T,W}^{WW}(q^2) \right],\end{aligned}$$

where  $\bar{\Pi}_{T,V}^{AB}(q^2) \equiv [\bar{\Pi}_T^{AB}(q^2) - \bar{\Pi}_T^{AB}(m_V^2)]/[q^2 - m_V^2]$  and the hat means the  $\overline{\text{MS}}$  coupling. In terms of the bar charges, the  $S$ ,  $T$  and  $U$  parameters can be written as

$$\begin{aligned}\frac{\bar{s}^2(m_Z^2)\bar{c}^2(m_Z^2)}{\bar{\alpha}(m_Z^2)} - \frac{4\pi}{\bar{g}_Z^2(0)} &= \frac{S}{4}, \\ \frac{\bar{s}^2(m_Z^2)}{\bar{\alpha}(m_Z^2)} - \frac{4\pi}{\bar{g}_W^2(0)} &= \frac{S+U}{4}, \\ 1 - \frac{\bar{g}_W^2(0)}{m_W^2} \frac{m_Z^2}{\bar{g}_Z^2(0)} &= \alpha T.\end{aligned}$$

## $S_Z$ and $T_Z$ (2)

The  $S$ ,  $T$  and  $U$  parameters:

$$\begin{aligned}\frac{\bar{s}^2(m_Z^2)\bar{c}^2(m_Z^2)}{\bar{\alpha}(m_Z^2)} - \frac{4\pi}{\bar{g}_Z^2(0)} &= \frac{S}{4}, \\ \frac{\bar{s}^2(m_Z^2)}{\bar{\alpha}(m_Z^2)} - \frac{4\pi}{\bar{g}_W^2(0)} &= \frac{S+U}{4}, \\ 1 - \frac{\bar{g}_W^2(0)}{m_W^2} \frac{m_Z^2}{\bar{g}_Z^2(0)} &= \alpha T.\end{aligned}$$

The last eq. can be written as

$$\frac{1}{\bar{g}_Z^2(0)} = \frac{1 - \alpha T + \bar{\delta}_G}{4\sqrt{2}G_F m_Z^2},$$

We are more interested in physics at the  $Z$  pole  $\implies$  Replace  $\bar{g}_Z(0)$  with  $\bar{g}_Z(m_Z^2)$ : ( $S_Z$  and  $T_Z$  parameters)

# $S_Z$ - $T_Z$ Plane Analysis

1. Calculate  $\mathcal{O}_i^{\text{th}}(\Delta S_Z, \Delta T_Z, \dots)$ , where  $\mathcal{O}_i$  are EW precision observables ( $\Gamma_Z, \sigma_h^0, A_f, \dots$ ).

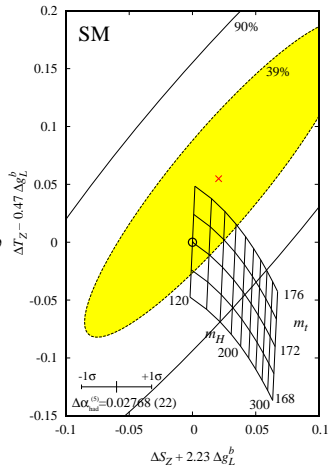
2. Construct the  $\chi^2$  function as

$$\chi^2 = \sum_{i,j} (\mathcal{O}_i^{\text{th}}(\Delta S_Z, \Delta T_Z, \dots) - \mathcal{O}_i^{\text{exp}}) \times (V^{-1})_{ij} (\mathcal{O}_j^{\text{th}}(\Delta S_Z, \Delta T_Z, \dots) - \mathcal{O}_j^{\text{exp}}),$$

where  $V$  is the covariance matrix,

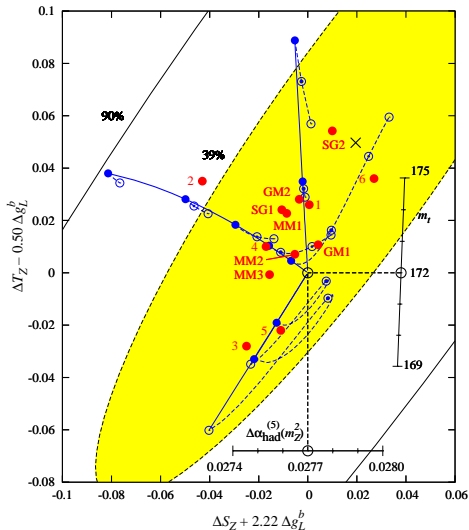
$$V_{ij} = (\delta \mathcal{O}_i^{\text{exp}})(\delta \mathcal{O}_j^{\text{exp}}) \rho_{ij}.$$

3. Find the minimum of  $\chi^2$  with respect to  $\Delta S_Z$ ,  $\Delta T_Z$  etc. Draw the contours  $\chi^2 - \chi^2_{\text{min}} = \text{const}$  if necessary.





# EW Precision Data vs MSSM, (II) $S_Z$ - $T_Z$ plane analysis



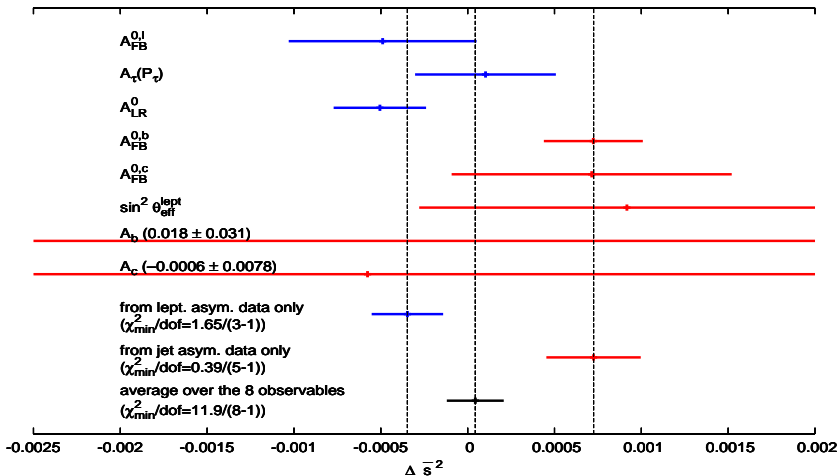
Using the final LEP EW precision data, we can give a constraint on MSSM contributions to  $S_Z$  and  $T_Z$ .

Our Results:

✓ All the sample points are within or close to the  $1\text{-}\sigma$  favored region.



# Problem in Jet Asymmetry Data?

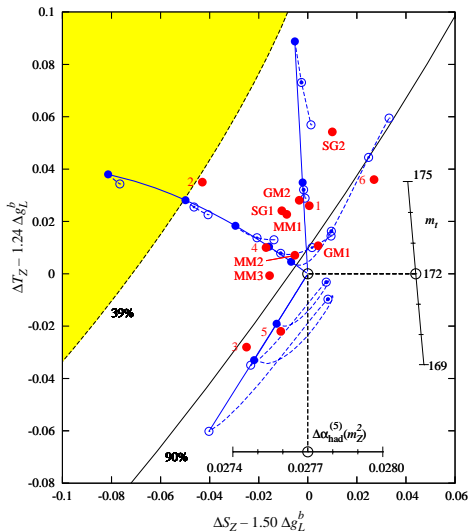


The value of the effective mixing angle  $\bar{s}^2$  determined from leptonic asymmetry data and that from jet asym. data do not agree very well

⇒ **problem in jet asym. data (or in the analysis)?**



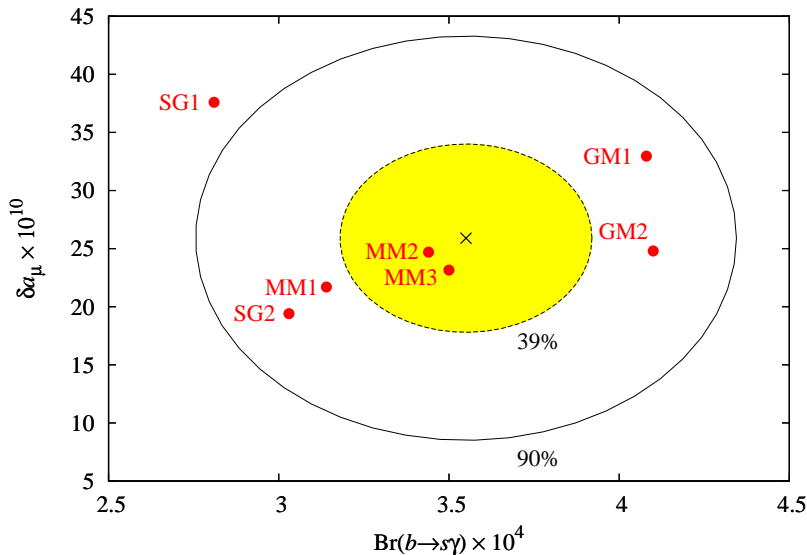
# EW Precision Data vs MSSM, (IV) fit w/o jet asymm. data



If we do not use the jet asymmetry data, the favored region shifts to the left. (Negative  $\Delta S_Z$  is favored.)

✓ Light sleptons are favored.

# Constraints from muon $g - 2$ and $b \rightarrow s\gamma$



# Summary

- We have studied the favored parameter region of MSSM using the results of the muon  $g - 2$  and the EW precision data.
- From muon  $g - 2$ : when  $\tan \beta = 10$ , the slepton mass of a few hundred GeV is favored. When  $\tan \beta = 50$ , the sleptons as heavy as 1 TeV are allowed within  $1-\sigma$ .
- From EW precision data: SUSY particles of a few hundred GeV are OK.
- In well-studied models like mSUGRA, Gauge Med. and Mirage Med. there still is some parameter region favored from muon  $g - 2$  and EW precision data.
- If we leave out the jet asymmetry data, light sleptons become more favored, which is favored from muon  $g - 2$  as well.